

Phys 2210 (15)

Solving eq'n of motion (N-2) generally means solving an "ORDINARY DIFFERENTIAL EQ'N", or ODE (Boas 8.1)  
 (Ordinary means no partial derivatives - later!)

Notation:  $y' = y'(x) = \frac{dy(x)}{dx} = \frac{dy}{dx}$ ;  $y'' = \frac{d^2y}{dx^2}$ , etc

$\dot{y} = \dot{y}(t) = \frac{dy(t)}{dt}$

Notation: "Order" of ODE = highest derivative

so e.g.  $m\ddot{r} = F(r)$  is a "2<sup>nd</sup> order ODE"

Notation: "Linear" ODE means  $y$  and all its deriv's appear without any power (or other functions)

Ex Linear ODEs:

$a y''(x) + c y(x) = 0$

$a(x) y''(x) + b(x) y'(x) = c(x)$

$\sqrt{x} y''(x) + \frac{\tan x}{x^3} y(x) = \ln x$

↑  
 Linear in y, fn(x) doesn't matter!

Ex: not Linear ODEs:

$(y''(x))^2 + y(x) = 0$

$y' \cdot y'' + y = 0$

not linear!  
 $y'' = -y^2$

NOTATION "Homogeneous" ODE means every term contains  $y$  and/or derivs of  $y$

Ex Homogeneous

$$ay''(x) + by'(x) + cy = 0$$

$$\sqrt{x}y''(x) - (x^2)y'(x) = 0$$

Not Homogeneous

$$ay'' + by' + cy = d$$

↑  
even just a constant!

Key ODE fact: Any linear ODE of order  $n$  has a general sol'n  $y(x)$  involving  $n$  independent constants which can only be determined by "boundary conditions" or "initial conditions" (like,  $y(x=0)$  or  $y'(x=0)$ , etc)

This fact is not always true for non-linear ODE's!

Example:  $y''(t) = -\omega^2 y(t)$ , we said the sol'n was

$$y = A \sin \omega t + B \cos \omega t$$

Two (2<sup>nd</sup> order!) constants

Notation: "Separable" ODE is when you can rearrange (treating "dx" as an algebraic object) so that

$$(\text{functions of } y) dy = (\text{functions of } x) dx$$

we like separable ODE's because you solve them by simply integrating both sides.

Example:  $\frac{dy}{dx} = f(x)$  is separable 1<sup>st</sup> order ODE

$$dy = f(x) dx \Rightarrow \int dy = \int f(x) dx$$

$$y = \int f(x) dx + C \rightarrow \text{our } \underline{\text{one}} \text{ constant!}$$

Example:  $m dv/dt = -cv^2$  ← Newton with drag, see soon!

∴  $\frac{dv}{v^2} = -\frac{c}{m} dt$  separated!  $f(v)$  on left,  $f(t)$  on right.

Integrate:  $-\frac{1}{v} = -\frac{c}{m} t + C_0$  ← our one constant, for our 1<sup>st</sup> order ODE!

You can find  $C_0$  if you have an "initial condition", e.g. if  $v(t=0) = v_0$  is given,  $-1/v_0 = 0 + C_0$ , + thus

$$v(t) = \frac{1}{\frac{c}{m} t - C_0} = \frac{1}{\frac{c}{m} t + \frac{1}{v_0}} = \frac{v_0}{1 + \frac{c v_0}{m} t}$$

2210

176

Example ("Math-y", not so physics-y". (Linear, 1<sup>st</sup> order ODE)

Solve  $xy' - y = xy$ , Given "boundary condition"  $y(x=1)=1$

$x \frac{dy}{dx} - y = xy$ . Let's separate x's + y's. Divide by x:

$\frac{dy}{dx} - \frac{y}{x} = y$ . Still not there! Group the y's:

$\frac{dy}{dx} = y/x + y = y(1 + \frac{1}{x})$  Ahhh! Now I see it!

$dy/y = (1 + \frac{1}{x}) dx$ . Separated! So, now just integrate

$\ln y = x + \ln x + C$  ← our one constant!

General sol'n:  $y = e^{(x + \ln x + C)} = x e^x e^C = \underbrace{C'}_x e^x$   
just rewrite the constant

[Boundary condition ("BC") tell us  $C'$ :  
 $y(x=1)=1 \Rightarrow 1 = C' \cdot 1 \cdot e^1$ , so  $y(x) = \frac{1}{e} x e^x = x e^{x-1}$ ]

2210 (18)

(Boas 8.3)

There are many tricks to solve ODE's.

Boas has a cool one for any linear, 1<sup>st</sup> order ODE,

$$y'(x) + P(x)y(x) = Q(x).$$

$\uparrow$   
any fn                       $\uparrow$   
any fn! (This makes it inhomogeneous!) )

Main idea:

1) Solve the homogeneous eq'n  $y_c'(x) + P(x)y_c(x) = 0$ , generally  
It's separable, so easy! This is the "complementary" eq'n.

2) Find some particular sol'n  $y_p$  to our diff eq, "y<sub>particular</sub>"

Then  $y_c + y_p = y_{\text{general}} + y_{\text{particular}}$  is the full general  
sol'n.

Boas finds a trick to find  $y_p$ , there's a formula, it's  
pretty elegant

→ Boas p. 401, eq 3.9.

Use it if you ever need it!

2210 (19)



Longmont United Hospital  
www.luhcares.org • 303.651.5111

1-2

Back to physics! (we discuss

$F = m\ddot{x}$  is a 2<sup>nd</sup> order

classical mechanics!)

= Spring Laptop

- Ivan, Jeffrey, Anthony  
Anich Ballard (Barbaro)

In phys 1110:  $m\ddot{x} = \text{const.}$

(gravity)

$$m\ddot{x} = -kx$$

(gs, SHM)

We can do more interesting & realistic problems now!

Including  $F = F(x)$ , or  $F(t)$ , or  $F(v)$ !

(often, writing as  $m\dot{v} = F$  will help)

1<sup>st</sup> class of new problems is PROJECTILES with DRAG

In real life,  $\vec{F} = \vec{F}_{\text{grav}} + \vec{F}_{\text{drag, or resistance}}$

← Taylor calls it  $\vec{f}$ .

$\vec{f}_{\text{drag}}$  depends (usually) on velocity, + points opposite  $\vec{v}$ ,

$$\vec{f} = -f(v) \hat{v}$$

$$\hat{v}$$

↳ Unit vector in  $\vec{v}$  direction

[Magnetism is  $\perp$  to  $\vec{v}$ , so this is a separate case]

$f(v) = |f|$  depends on speed  $v = |\vec{v}|$ .

$f(v)$  could be anything! Taylor's theorem says (soon!)

$f(v) \approx a + bv + cv^2 + \dots$  (Let's look at leading terms)

Phys 1110 considered  $F_{\text{friction}} = \mu N = \text{constant}$ , indep. of  $|\vec{v}|$ , but this is not considered "drag". So, let's set  $f(v=0) = 0$ ,

$$f(v) = bV + cV^2$$

Linear  $\leftarrow$  quadratic drag.

Low speeds, dominates.  $\uparrow$  High speeds, this dominates

- The physics of  $f_{\text{linear}}$  is viscosity. (Move through mud, or molasses, or even air)

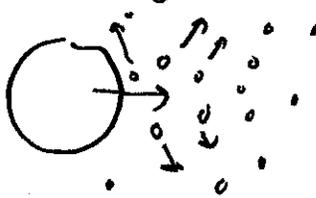
viscous friction



So "b" depends on size of object

and some property (viscosity) of medium

- The physics of  $f_{\text{quadratic}}$  is from "moving the medium out of the way"



you must accelerate medium particles away.

the faster you go, the faster the recoiling particles must bounce to get away

and you encounter more particles each second

Both of these involve a factor of "v", hence  $v^2$ .

So "c" depends on area of object (how much it "sweeps out")

and also density,  $\rho$ , of the medium (more molecules to sweep)

2210 (21)

For spherical objects, then, of diameter  $D$

$$f_{\text{linear}} = bV = (\beta D) V \quad \text{this is "Stokes Law"}$$

$\downarrow$   $\downarrow$   
 measure of viscosity      Diameter

$$f_{\text{quad}} = cV^2 = (\gamma D^2) V^2$$

$\downarrow$  Area goes like  $D^2$ .  
 $\downarrow$  depends on density  $\rho$ , also shape/streamlining

$$\equiv \left( \frac{1}{2} C_D \rho A \right) V^2 \quad \text{this defines } C_D$$

$\downarrow$  cross sectional area =  $\pi D^2/4$   
 $\downarrow$  Density of medium  
 $\downarrow$  "the drag coefficient"

$C_D$  is dimensionless

Depends on shape, mostly

(convince yourself!)  $\rightarrow (\rho \cdot A \cdot V^2 = \frac{\text{kg}}{\text{m}^3} \cdot \text{m}^2 \cdot \frac{\text{m}^2}{\text{s}^2} = \frac{\text{kg m}}{\text{s}^2})$

$\downarrow$  (Air has  $\rho \approx 1.29 \text{ kg/m}^3$ )

$C_D$  (Hummer)  $\approx .57$

$C_D$  (sphere) = .5

$C_D$  (Prius) = .26

$C_D$  (747) = .03

$$\frac{f_{\text{quad}}}{f_{\text{lin}}} = \frac{cV^2}{bV} = \frac{\gamma D_{\text{diam}}}{\beta} V = \left( \frac{1}{2} \frac{C_D \rho}{\beta} \frac{\text{Area}}{\text{Diam}} \right) V \propto \text{"Reynolds \#"}$$

For Big  $V$ ,  $f_{\text{quad}}$  dominates; this sets the scale!

2210 -

(22)

Terminal velocity: As  $v \uparrow$ ,  $f$  also  $\uparrow$ , so as objects fall, eventually  $f$  balances weight, + acceleration  $\rightarrow 0$ . That means constant  $v = "V_{\text{TERMINAL}}"$ .

$\sim 120 \text{ mi/hr} \approx 55 \text{ m/s}$  for skydiver in freefall, (but depends on body orientation, clothes, etc)

Linear drag:  $\vec{f}(v) = -b \underbrace{v \hat{v}}_{\vec{v}} = -b \vec{v}$

(so if  $F_{\text{net}} = 0 \Rightarrow \underline{b V_{\text{term}}} = mg$ )

or  $\boxed{V_{\text{term}} = \frac{mg}{b}} = \frac{mg}{\beta D}$

Checks: Lower  $\beta \Rightarrow$  higher  $V_T$  ✓  
 Bigger objects  $\Rightarrow$  lower  $V_T$ , ✓  
 elephants die, ants survive!  
 Denser objects of given size ✓  
 $\Rightarrow$  higher  $V_T$ , I believe!

Quadratic drag:  $\vec{f}(v) = -c v^2 \hat{v} = -c |v| \vec{v}$

So if  $F_{\text{net}} = 0 \Rightarrow c V_T^2 = mg$

or  $\boxed{V_{\text{term}} = \sqrt{\frac{mg}{c}}} = \sqrt{\frac{mg}{\frac{1}{2} \rho c_D A}}$

(But, this is just "steady state" ...)

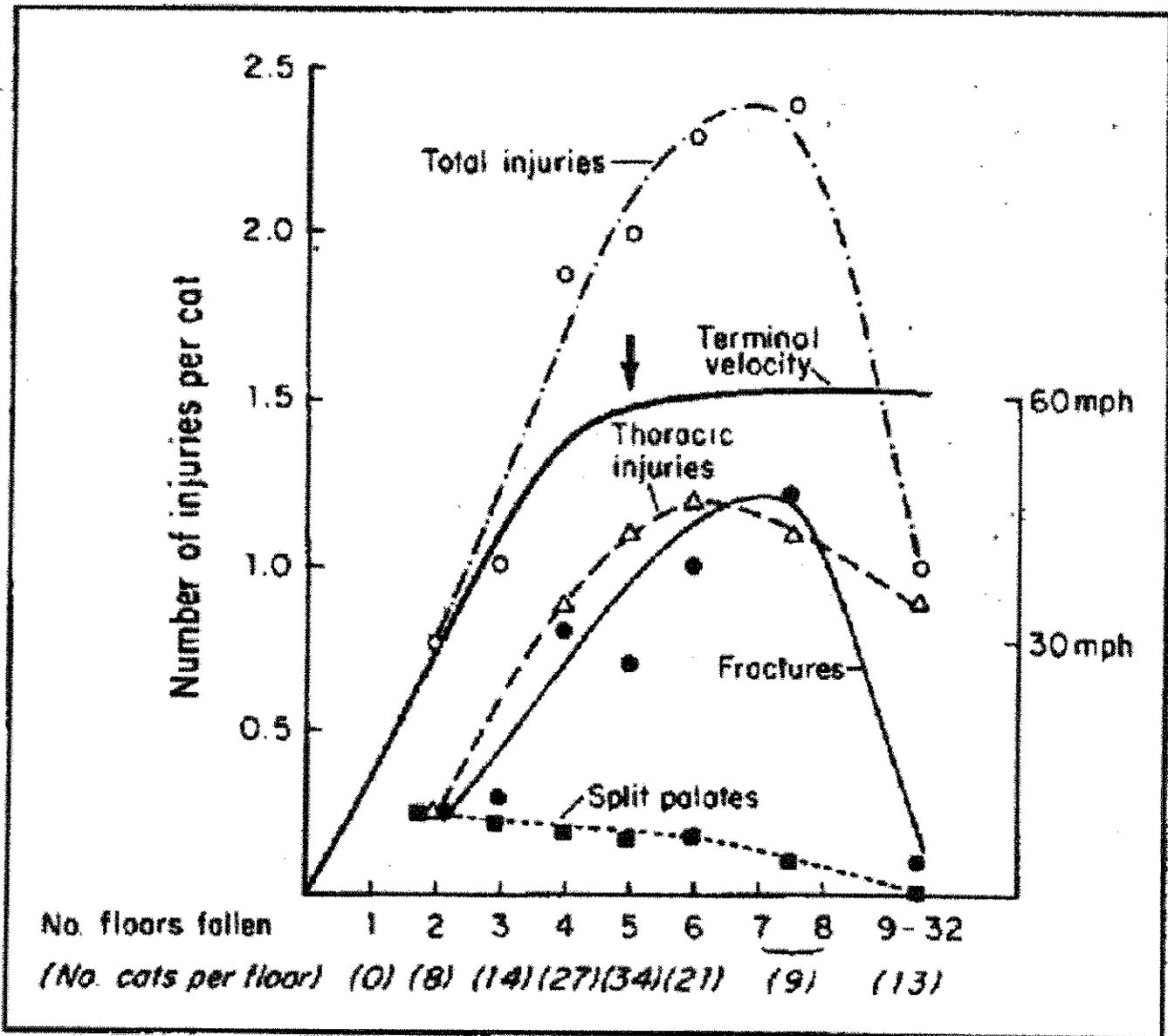


Figure 2—Relationship of injuries to distance fallen and velocity in 132 cats with high-rise syndrome: ↓ points to terminal velocity (—); total number of injuries/cat (○, - - - -); number of thoracic injuries (pulmonary contusions + pneumothorax)/cat (△, - - -); number of fractures/cat (●, —); number of split palates/cat (■, - - - -).

2210 226

Motion with drag in 1-D.

$$\vec{F}_{\text{net}} = m \vec{a} \quad \text{says} \quad m \dot{\vec{v}} = -b \vec{v}, \quad \text{so} \quad m \dot{v}_x = -b v_x$$

thus,  $\dot{v}_x = -\frac{b}{m} v_x$ . True if moving right ( $v_x > 0$ ) or left ( $v_x < 0$ ), convince yourself!

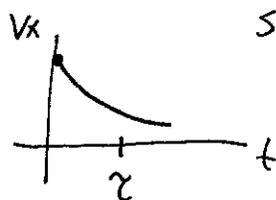
Separable!  $\frac{dv_x}{v_x} = -\frac{b}{m} dt$ . Integrate  $\Rightarrow \ln v_x = -\frac{b}{m} t + C$

If  $v_x(t=0) \equiv v_{x0}$  is given, then  $C = \ln v_{x0}$ ,  $v_x = v_{x0} e^{-bt/m}$

Note:  $[b/m] = \left[ \frac{N/(m/s)}{kg} \right] = \frac{kg \cdot m \cdot s}{s^2 \cdot m \cdot kg} = sec^{-1}$ . So, we define

a natural time  $\tau \equiv m/b$ !

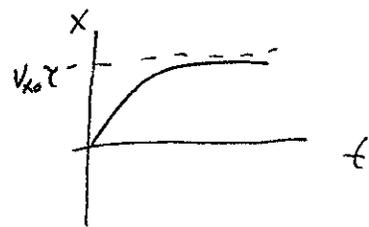
$$v_x(t) = v_{x0} e^{-t/\tau}$$



Slows down in "typical time"  $\tau$ .

To get  $x(t)$ ,  $\frac{dx}{dt} = v_x$ , so  $x(t) - x(0) = \int_{t=0}^t v_x dt = v_{x0} (-\tau) e^{-t/\tau} \Big|_0^t$

$$\text{so } \underline{x(t) - x(0) = -v_{x0} \tau (e^{-t/\tau} - 1)}$$



Note: As  $t \rightarrow \infty$ ,  $x(t) - x(0) \approx v_{x0} \tau$  is finite

It "grinds to a halt", (the same distance it would have drifted without drag in the natural time  $\tau$ .)

Vertical

# Motion, with drag, arbitrary initial conditions

Start in 1-D (e.g. dropping) My convention,  $+y \uparrow$   $\uparrow bv$   
 not Taylor's  $\downarrow mg$   
 (up = "plus")

Falling object:

$$\vec{F}_{net} = m \vec{a} \Rightarrow mg(-\hat{y}) - b\vec{v} = m \dot{\vec{v}}_0 \quad (\text{Taylor's})$$

In 1-D, ~~the same as above~~  $m \dot{v}_y = -mg - b v_y$  ~~no longer valid!~~

Separable! (or, use Boas trick, try that yourself) Boas p.40

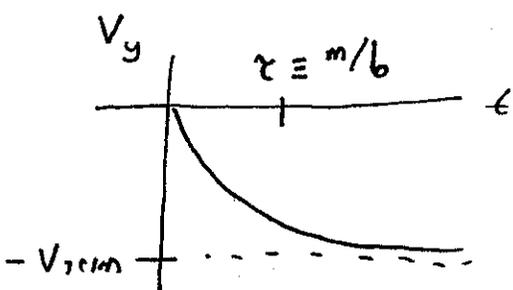
$$\frac{dv_y}{g + \frac{b}{m} v_y} = -dt \Rightarrow \frac{m}{b} \ln\left(g + \frac{b v_y}{m}\right) = -t + C$$

$$\text{so } g + \frac{b v_y}{m} = A e^{-bt/m}$$

Note: as  $t \rightarrow \infty$ ,  $v_y \rightarrow -g \frac{m}{b}$ , (see previous page!)  $\uparrow$   
 $= -v_{term}$  (sign is correct with my convention.)

If start from rest,  $v_y(t=0) = 0 \Rightarrow A = g$ , and so

$$v_y(t) = \frac{mg}{b} \left( e^{-bt/m} - 1 \right) = -v_{term} (1 - e^{-bt/m})$$



[Note: small  $t \Rightarrow e^{-bt/m} \approx 1 - bt/m + \dots$   
 $v_y \approx -v_{term} \frac{bt}{m} \approx -gt$   
 Ah, same as no drag, for small  $t$ ]

2210 -24

If we start with non-zero  $v_{0y}$ , then

$$g + \frac{b v_y(t=0)}{m} = A, \text{ so } \left( \text{using } v_t \equiv \frac{mg}{b}, \text{ and } \tau \equiv \frac{m}{b} \right)$$

$$v_y(t) = \frac{m}{b} \left( -g + A e^{-bt/m} \right) = -v_t (1 - e^{-t/\tau}) + \underbrace{v_{0y} e^{-t/\tau}}_{\text{extra!}}$$

If want motion, integrate  $v_y = dy/dt$ ,

$$\text{so } y(t) - y(0) = \int_0^t v_y \cdot dt = \int_0^t \left[ -v_t + (v_t + v_{0y}) e^{-t/\tau} \right] dt$$

$$= -v_t \cdot t - \tau (v_t + v_{0y}) (e^{-t/\tau} - 1)$$

(See Taylor for sketches...)

Comment: Note that  $\tau = m/b$  is a "natural time constant" here.

It has units of  $\text{kg}/(\text{N/m/s}) = \frac{\text{kg}}{\frac{\text{kg} \cdot \text{m}}{\text{s}^2} \cdot \frac{\text{s}}{\text{m}}} = \text{sec} \checkmark$

(It tells you the natural "decay time" for this motion, the timescale for speed to approach terminal velocity)

Comment: If  $v_y(0)$  is +, my sol'n's + eq'n's all still work.

So, ball could initial be thrown up (or released, or thrown down)

The above is still all fine.

2210 24b.

What about Quadratic drag,  $f_D \sim cv^2$ ?

↳ Consider pure 1-D motion with drag (only)

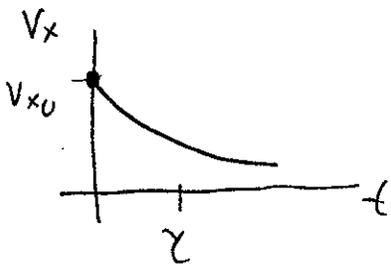
$$\vec{F} = m\vec{a} \text{ says } m\vec{\dot{v}} = -c|\vec{v}|\vec{v}, \text{ so } m\dot{v}_x = -c|v_x|v_x$$

If  $v_{x0} > 0$ , (then  $|v_x| = v_x > 0$ ),  $\dot{v}_x = -\frac{c}{m}v_x^2$ . Separable!

$$\frac{dv_x}{v_x^2} = -\frac{c}{m} dt \Rightarrow -\frac{1}{v_x} = -\frac{c}{m}t + C$$

As usual, if  $v_{x0}$  is given at  $t=0$ , 
$$v_x(t) = \frac{1}{\frac{ct}{m} + \frac{1}{v_{x0}}} = \frac{v_{x0}}{1 + \frac{cv_{x0}t}{m}}$$

Here, we have a new "natural time",  $\tau \equiv \frac{m}{cv_{x0}}$ . (check units yourself!)



Similar to linear drag in shape, but inverse time rather than exponential decay.

For  $x(t)$ , use  $\frac{dx}{dt} = v_x(t)$  and integrate:

$$\begin{aligned} x(t) - x(0) &= \int_{t=0}^t \frac{v_{x0}}{1 + \frac{cv_{x0}}{m}t} dt = v_{x0} \tau \ln(1 + t/\tau) \Big|_0^t \\ &= \frac{m}{c} \ln(1 + t/\tau) \end{aligned}$$

Note: At  $t \rightarrow \infty$ ,  $x$  does continue to grow, but  $\ln(1 + t/\tau)$

is a very slow "blowup". ( $\ln$  is a "boring function"?)

This is different from linear drag, where  $x(\infty)$  was finite.

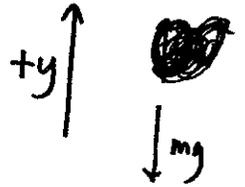
2210 (25)

In vertical motion (with gravity),

What about QUADRATIC DRAG? (Recall  $V_t^2 = \frac{mg}{c}$ )

$$F = ma \Rightarrow m \dot{v}_y = -mg - c|v_y|v_y.$$

$$\text{so } \dot{v}_y = -g - \frac{c}{m}|v_y|v_y = -g \left(1 + \frac{|v_y|v_y}{V_t^2}\right)$$



If drop from rest,  $v_y$  is always negative,  $|v_y| = -v_y$ , and

falling:  $\dot{v}_y = -g \left(1 - \frac{v_y^2}{V_t^2}\right)$  1<sup>st</sup> order, non-linear, separable.

$$\frac{dv_y}{1 - v_y^2/V_t^2} = -g dt \quad \text{Can integrate!}$$

Look it up, or just do it by integration by parts:

$$\text{Trick: } \frac{1}{(1-x)(1+x)} = \frac{1/2}{1-x} + \frac{1/2}{1+x}$$

convince yourself!

$$\text{Let } x = v_y/V_t \quad dx = dv_y/V_t$$

$$\int \frac{dv_y}{1 - v_y^2/V_t^2} = \int -g dt = -g t + C$$

$$V_t \int \frac{dx}{(1-x)(1+x)} = -\frac{1}{2} V_t \ln(1-x) + \frac{1}{2} V_t \ln(1+x)$$

$$\text{So, } \frac{V_t}{2} \ln\left(\frac{1+v_y/V_t}{1-v_y/V_t}\right) = -g t + C$$

If start at  $v_y=0$ ,  
 $C=0$ ,  
 convince yourself!

2210

(26)

So falling from rest, in quadratic drag case:

$$\frac{1 + v_r/v_t}{1 - v_r/v_t} = e^{-2gt/v_t}, \text{ so (convince yourself!)}$$

$$\frac{v_r}{v_t} = \frac{e^{-2gt/v_t} - 1}{e^{-2gt/v_t} + 1}$$

Note,  $t \rightarrow \infty$ ,  $\frac{v_r}{v_t} \rightarrow -1$ ,  
that's my sign convention!

Review of Hyperbolic functions (Boas ch 2.12)

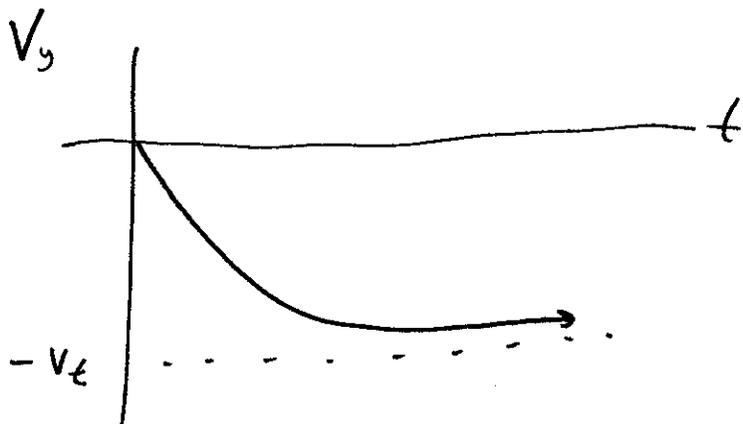
$$\sinh x = \frac{1}{2} (e^x - e^{-x})$$

$$\cosh x = \frac{1}{2} (e^x + e^{-x})$$

$$\tanh x = \frac{\sinh x}{\cosh x} = \frac{e^x - e^{-x}}{e^x + e^{-x}} = \frac{e^{2x} - 1}{e^{2x} + 1}$$

$$\text{So } \frac{v_r}{v_t} = \tanh\left(-\frac{gt}{v_t}\right) = -\tanh\frac{gt}{v_t}$$

↳ Because of my sign convention



Qualitatively similar  
to linear drag case.

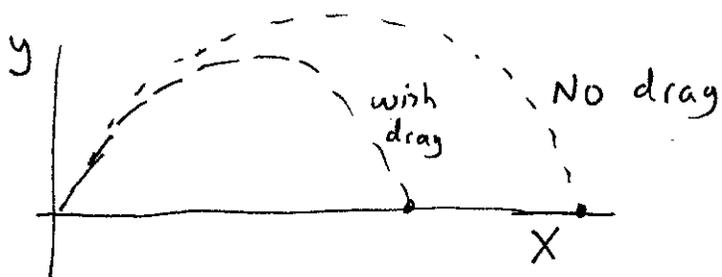
2210 - 27

(Taylor 2.3)

Now, at last, 2-D motion, "Projectile Motion"

Let's start w. linear drag:

$$m\vec{a} = \vec{F}$$



$$\text{so } \begin{cases} m \dot{v}_x = -b v_x \\ m \dot{v}_y = -mg - b v_y \end{cases}$$

Recall our definitions  $\left\{ \begin{array}{l} \tau = m/b \text{ "natural" time} \\ v_T = mg/b \text{ "terminal" speed} \end{array} \right.$

We've solved both of the eq'ns of motion

$$x(t) = \tau v_{x0} (1 - e^{-t/\tau}) \quad \leftarrow \text{Notes p. 22b, with } x(0) = 0$$

$$y(t) = -v_{Tt} + \tau (v_{Tt} + v_{y0}) (1 - e^{-t/\tau}) \quad \leftarrow \text{p. 24, with } y(0) = 0$$

Can find  $y(x)$  by eliminating time in 1<sup>st</sup> eq'n:

$$e^{-t/\tau} = 1 - x/v_{x0}\tau \quad \text{or} \quad t = -\tau \ln(1 - x/v_{x0}\tau)$$

Thus

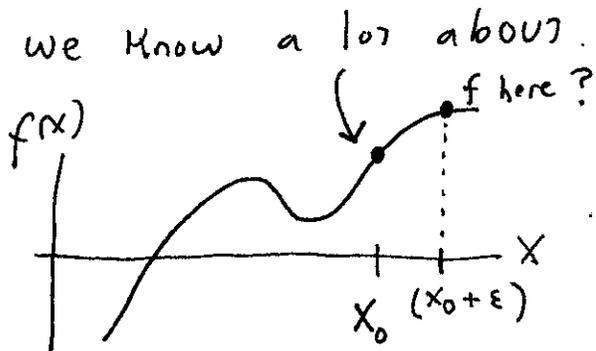
$$y(x) = +v_{T}\tau \ln\left(1 - \frac{x}{v_{x0}\tau}\right) + \frac{v_{Tt} + v_{y0}}{v_{x0}} x.$$

Not the prettiest of eq'ns!

For small drag (small  $b \Rightarrow$  long  $\tau$ ) we can approximate this formula very nicely. This involves one of the most useful approximation approaches, used throughout physics. So, a brief math interlude!

## Taylor Series expansions (approximations)

we often need to know  $f(x_0 + \epsilon)$ , a function near a place  
 we know a lot about.



(E.g. previous page, we want  
 $\ln(1 - \frac{x}{v_{x_0 c}})$  for small  $\frac{x}{v_{x_0 c}}$ )

This is one of the premier Math tools/tricks!

Consider  $f(x)$  near  $x_0$ . Most functions can be "written as a

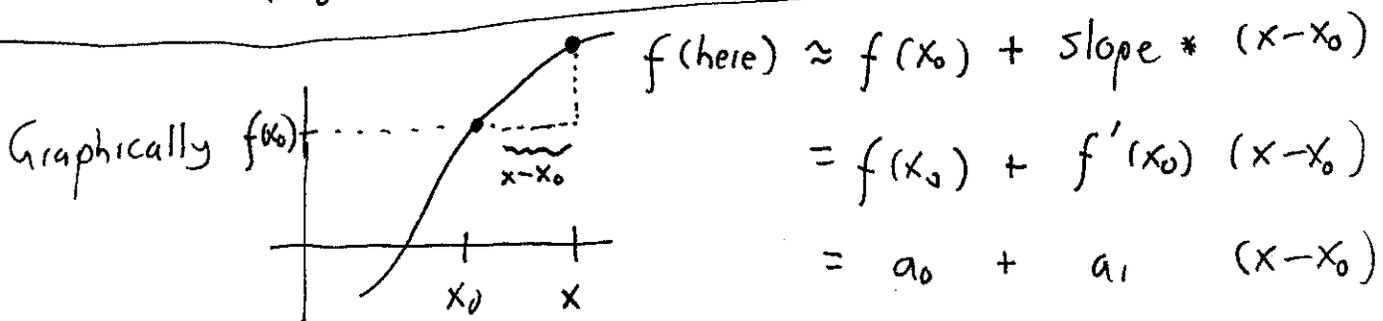
power series

$$f(x) = a_0 + a_1 (x-x_0) + a_2 (x-x_0)^2 + \dots$$

"linear"
"quadratic"

↓
↓
↓
constant

$$f(x) = \sum_{n=0}^{\infty} a_n (x-x_0)^n$$



(This would be the 1<sup>st</sup> term in that Taylor series  $\uparrow$ )  
 (treating  $f(x)$  as approximately linear near  $x_0$ .)

2210 -29.

Notice  $f(x_0) = a_0 + a_1(x_0 - x_0) + a_2(x_0 - x_0)^2 + \dots$   
 $= a_0$

and  $f'(x) = 0 + a_1 + 2 \cdot a_2(x - x_0) + 3 \cdot a_3(x - x_0)^2 + \dots$

so  $f'(x_0) = \cancel{0} a_1 + 0 + 0 + \dots$

and  $f''(x) = 0 + 0 + 2 \cdot 1 a_2 + 3 \cdot 2 a_3(x - x_0) + \dots$

so  $f''(x_0) = 2 a_2 + 0 + 0 + \dots$

See the pattern?  $a_0 = f(x_0)$

$$a_1 = f'(x_0)$$

$$a_2 = f''(x_0) / 2$$

$$a_3 = f'''(x_0) / 3!$$

$$\dots a_n = f^{(n)}(x_0) / n!$$

So our "Taylor series around the point  $x_0$ " is

$$f(x) = f(x_0) + f'(x_0)(x - x_0) + \frac{f''(x_0)}{2!}(x - x_0)^2 + \frac{f'''(x_0)}{3!}(x - x_0)^3 + \dots$$

(Notation: If  $x_0 = 0$ , we call this a Maclaurin series)

If  $x$  is near  $x_0$ , so  $x - x_0 \equiv \epsilon$  is small

$$f(x_0 + \epsilon) \approx f(x_0) + f'(x_0) \cdot \epsilon + \frac{f''(x_0)}{2!} \epsilon^2 + \dots$$

(Also the Taylor series)

2210 -30.

Example  $f(x) = e^x$ , and  $x_0 = 0$ .

so  $f'(x) = e^x$ ,  $f''(x) = e^x$ , etc, and at  $x_0 = 0$ , all  $= e^0 = 1$ .

$\therefore$  Maclaurin series for  $e^x$  is

$$e^x = f(0) + f'(0)(x-0) + \frac{f''(0)}{2!}(x-0)^2 + \dots$$
$$= 1 + x + \frac{x^2}{2!} + \frac{x^3}{3!} + \dots$$

Example  $f(x) = \sin(x)$ , and  $x_0 = 0$ .

$f'(x) = \cos(x)$ ,  $f''(x) = -\sin(x)$ ,  $f'''(x) = -\cos(x)$ ,  $f^{(4)}(x) = \sin(x)$   
 $f^{(0)}(0) = 0$ ,  $f'(0) = 1$ ,  $f''(0) = 0$ ,  $f^{(3)}(0) = -1$ ,  $f^{(4)}(0) = 0$ , etc.

$$\sin x = f(0) + f'(0)x + \frac{f''(0)}{2!}x^2 + \dots$$

$$\sin x = x - \frac{x^3}{3!} + \frac{x^5}{5!} - \dots$$

See Front Flyleaf of Taylor for the common ones.

Memorize 'em, you'll use them the rest of your life!

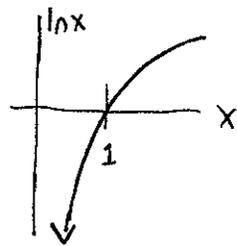
( $\sin x$ ,  $\cos x$ ,  $e^x$ , ...) If  $x$  small, 1<sup>st</sup> term or 2 may be all you need!!

$$\sin x \approx x - \frac{x^3}{6} + \dots \Rightarrow \sin(.1 \text{ rad}) = .1 - .001/6 = \underline{\underline{.0998}}$$

$\uparrow$   
"Leading term", or "first order term"

2210 -31

Example  $f(x) = \ln x$ . This blows up at  $x=0$ ,  
 so Maclaurin series at  $x_0=0$  is useless. BUT,  
 Taylor series at  $x_0=1$  is handy!



$$f(x) = \ln x \quad \Rightarrow \quad f(x_0) = \ln 1 = 0$$

$$f'(x) = 1/x = x^{-1} \quad \Rightarrow \quad f'(x_0) = 1^{-1} = 1$$

$$f''(x) = (-1)x^{-2} \quad \Rightarrow \quad f''(x_0) = (-1)(1)^{-2} = -1$$

$$f'''(x) = (-1)(-2)x^{-3} \quad \Rightarrow \quad f'''(x_0) = (+2!)(1)^{-3} = +2!$$

$$\begin{aligned} \text{So } f(x) &= f(x_0) + f'(x_0)(x-1) + \frac{f''(x_0)}{2!}(x-1)^2 + \dots \\ &= 0 + 1(x-1) - \frac{(x-1)^2}{2} + \frac{2!}{3!}(x-1)^3 - \frac{3!}{4!}(x-1)^4 + \dots \end{aligned}$$

$$\text{So } \ln x = (x-1) - \frac{(x-1)^2}{2} + \frac{(x-1)^3}{3} - \frac{(x-1)^4}{4} \quad \leftarrow \text{NOT factorials!!}$$

I prefer to rewrite this for  $x-1 = \epsilon$ , or  $x = 1 + \epsilon$

$$(\text{Near } x=1) \quad \ln(1+\epsilon) = \epsilon - \frac{\epsilon^2}{2} + \frac{\epsilon^3}{3} - \dots$$

$$\text{or, (convince yourself)} \quad \ln(1-\epsilon) = -(\epsilon + \frac{\epsilon^2}{2} + \frac{\epsilon^3}{3} + \dots)$$

(Makes sense that this is negative, look @ the graph!)  $\uparrow$

2210 -32.

one last example, a very handy one.

$f(x) = X^a$  where  $a$  might not be an integer.

(Around  $x=0$ , you already have a series... with 1 term!)

But if  $x_0=1$ , you get the wonderfully useful "Binomial" Series

$$f'(x) = a x^{a-1} \quad \longrightarrow \quad f'(x_0=1) = a$$

$$f''(x) = a(a-1)x^{a-2} \quad \longrightarrow \quad f''(x_0=1) = a(a-1)$$

$\therefore$  If  $a$  is an integer, the series is finite, but in general

$$f(x) = f(x_0) + (x-x_0) f'(x_0) + \frac{(x-x_0)^2}{2!} f''(x_0) + \dots$$

$$= 1 + (x-x_0) a + \frac{(x-x_0)^2}{2!} a(a-1) + \dots$$

Once again, think of  $x = x_0 + \epsilon = 1 + \epsilon$

so  $x-x_0 = \epsilon$  is small, and

$$(1+\epsilon)^a \approx 1 + a\epsilon + \frac{a(a-1)}{2!} \epsilon^2 + \frac{a(a-1)(a-2)}{3!} \epsilon^3 + \dots$$

$$\text{e.g. } \sqrt{1+\epsilon} = (1+\epsilon)^{1/2} \approx 1 + \frac{1}{2}\epsilon + \frac{1}{2!} \left(\frac{1}{2}\right)\left(-\frac{1}{2}\right)\epsilon^2 + \dots$$

(If you "truncate", it's approximate, but accurate if  $\epsilon \ll 1$ .)

Polynomials are much simpler to deal with. It's rare in real life that physics/engineering problems can be solved exactly, but it's common they can be approximated by a series (which is as accurate as you need, if you take enough terms)

Notation: Analytic functions have an  $\infty$  # of finite derivatives at every point, and thus the Taylor series is "exact".

Back to our 2-D projectile with ~~linear~~ <sup>linear</sup> drag,

$$y(x) = v_{0y} \tau \ln(1 - x/v_{x0} \tau) + \left( \frac{v_{0x} + v_{y0}}{v_{x0}} \right) x \quad \leftarrow \text{Exact.}$$

If drag is small (or if  $x$  is small)  $\frac{x}{v_{x0} \tau} \equiv \epsilon$  is small,

$$\text{and } \ln(1 - \epsilon) \approx -(\epsilon + \epsilon^2/2 + \epsilon^3/3 + \dots)$$

$$\text{so } y(x) \approx +v_{0y} \tau \left( \frac{x}{v_{x0} \tau} + \left( \frac{x}{v_{x0} \tau} \right)^2 \cdot \frac{1}{2} + \left( \frac{x}{v_{x0} \tau} \right)^3 \cdot \frac{1}{3} + \dots \right) + \left( \frac{v_{0x}}{v_{x0}} x + \frac{v_{y0}}{v_{x0}} x \right)$$

$$y(x) = \frac{v_{y0}}{v_{x0}} x - \frac{1}{2} \frac{v_{0y}}{v_{x0}^2} x^2 - \frac{1}{3} \frac{v_{0y}}{v_{x0}^3} x^3 + \dots$$

↑  
Cancels with very 1st term!!

(Let's pause for a sec + figure out  $y(x)$  when there is no drag)

oh, also note  $v_{0y}/\tau = \frac{mg/b}{m/b} = g$ , so 2<sup>nd</sup> term is  $-\frac{1}{2} \frac{g x^2}{v_{x0}^2}$

2210 - 34

In vacuum:  $x = v_{x0} t$ ← use this to find  $t$ 

$$y = v_{y0} t - \frac{1}{2} g t^2$$

← plug in here, + get

$$y(x)_{\text{no drag}} = \frac{v_{y0}}{v_{x0}} x - \frac{1}{2} \frac{g}{v_{x0}^2} x^2$$

that's it, a parabola!

$$y(x)_{\text{with drag}} = \frac{v_{y0}}{v_{x0}} x - \frac{1}{2} \frac{g}{v_{x0}^2} x^2 - \frac{1}{3} \frac{v_t}{v_{x0}^3} x^3 + \dots$$

Sweet! The 2 leading terms match, so at 1st (small  $x$ ), the 2 curves look identical. It's the 3<sup>rd</sup> order term that slowly but surely "deviates" the path, and is negative so our drag trajectory is (of course!) below the ideal path.

How far do we go? The range  $R$ this is  $x$ , for which  $y(x=R) = 0$ .

For no drag,  $0 = \frac{v_{y0}}{v_{x0}} R_{\text{vac}} - \frac{1}{2} g \frac{R_{\text{vac}}^2}{v_{x0}^2}$ . of course,  $R=0$  works, but also

$$R_{\text{vac}} = \frac{2 v_{y0} v_{x0}}{g}$$

Phys 1110 formula!

2210 - 35

With drag,  $V_t \tau \ln(1 - R/V_{x0}\tau) + \frac{V_t + V_{y0}}{V_{x0}} R = 0$

If  $R=0$ , this is satisfied, but there's a 2<sup>nd</sup> sol'n.

In general, you can't solve this eq'n analytically for it.

But if  $R/V_{x0}\tau \equiv \epsilon \ll 1$ , we can approximate it:

$$y^{(R)}_{\text{with drag}} = \frac{V_{y0}}{V_{x0}} R - \frac{1}{2} \frac{g}{V_{x0}^2} R^2 - \frac{1}{3} \frac{V_t}{V_{x0}^3 \tau^2} R^3 + \dots \text{(terms of order } R^4 \text{)}$$

Setting this to 0, and dividing out one power of  $R$  gives

$$0 = \frac{V_{y0}}{V_{x0}} - \frac{1}{2} \frac{g}{V_{x0}^2} R - \frac{1}{3} \frac{V_t}{V_{x0}^3 \tau^2} R^2 + \text{Order}(R^3)$$

$$\text{or } R = \frac{2V_{y0}V_{x0}}{g} - \frac{2}{3} \frac{V_t}{V_{x0}g\tau^2} R^2 + \mathcal{O}(R^3)$$

$$R = \text{~~0~~} R_{\text{vac}} - \frac{2}{3} \frac{V_t}{V_{x0}g\tau^2} R^2 + \mathcal{O}(R^3)$$

↑  
range is reduced.

Method 1: If  $R \approx R_{\text{vac}}$ , then plug that in to approximate  $R$

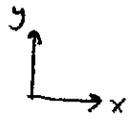
Method 2 (better, formally): Solve this quadratic eq'n for  $R$ !

Then, if you like, use  $\sqrt{1+\epsilon} \approx 1 + \frac{1}{2}\epsilon$  to simplify, where

$\epsilon$  here will be some small term...

2210 -36

# Quadratic drag for Projectile Motion



$$m \dot{\vec{v}} = m \vec{g} - c |\vec{v}| \vec{v}, \text{ so } m \dot{v}_x = -c \sqrt{v_x^2 + v_y^2} v_x$$

$$m \dot{v}_y = -mg - c \sqrt{v_x^2 + v_y^2} v_y$$

This is nasty - it's a "coupled" and non-linear equations. I can't solve it, I don't know any tricks. But, there is a practical way: If you know "initial conditions", i.e.  $v_x(t_0)$  and  $v_y(t_0)$

then e.g.  $m \dot{v}_x = -c \sqrt{v_x^2 + v_y^2} v_x$  means

$$m \frac{v_x(t_0 + \Delta t) - v_x(t_0)}{\Delta t} \approx -c \sqrt{v_x^2(t_0) + v_y^2(t_0)} v_x(t_0)$$

you know everything but  $v_x(t_0 + \Delta t)$ . Simple, it's linear, find it!

Similarly, find  $v_y(t_0 + \Delta t) \approx v_y(t_0) - \Delta t g - \frac{\Delta t c}{m} \sqrt{v_x^2(t_0) + v_y^2(t_0)} v_y(t_0)$

then, iterate, find  $v_x(t_0 + 2\Delta t)$ , etc. It's a numerical approach

This is what "ND Solve" does in MMA!

you get  $v_x(\text{any } t)$  and  $v_y(\text{any } t)$ , as numbers.

And thus  $\frac{dx}{dt} = v_x(t)$  can similarly give you  $x(\text{any } t)$ .

This might be the engineer's approach, and it's e.g. what the PhET sim does to graph  $y(x)$ .

2210 -366

This wraps up Ch. 2:

We're punting "Motion in Uniform B"

(i.e.  $\vec{F} = q \vec{v} \times \vec{B}$ ), at least for now. (You'll

cover this explicitly in Phys 3310, !)